Parallel Computing with Adaptive Mesh Refinement Cosmological simulations with RAMSES

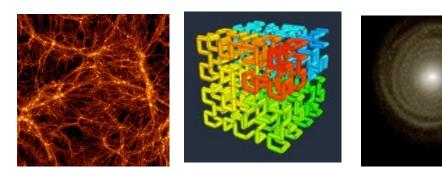
Romain Teyssier, Markus Wetzstein



University of Zurich

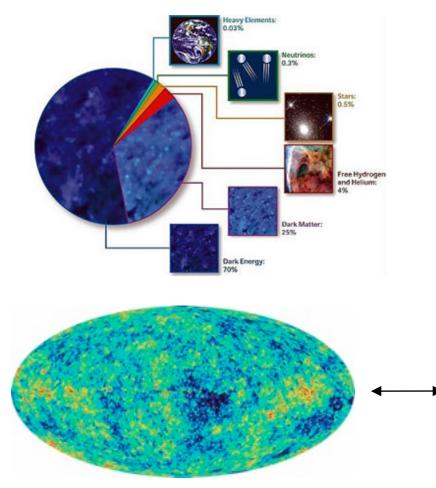


HP2C High Performance and High Productivity Computing

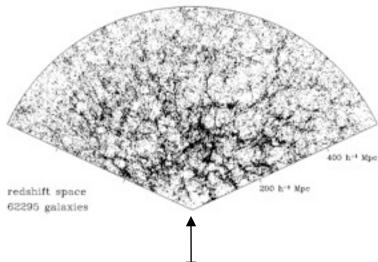


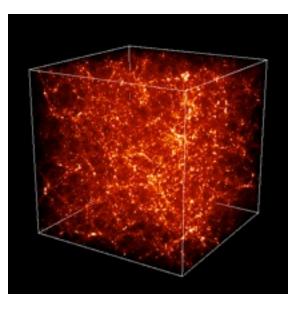
- Cosmological simulations
- Adaptive Mesh Refinement techniques
- Graded octree data structure
- Hydrodynamics with AMR
- Adaptive time stepping
- Parallel computing with RAMSES
- Cosmological simulations with RAMSES

Cosmological simulations

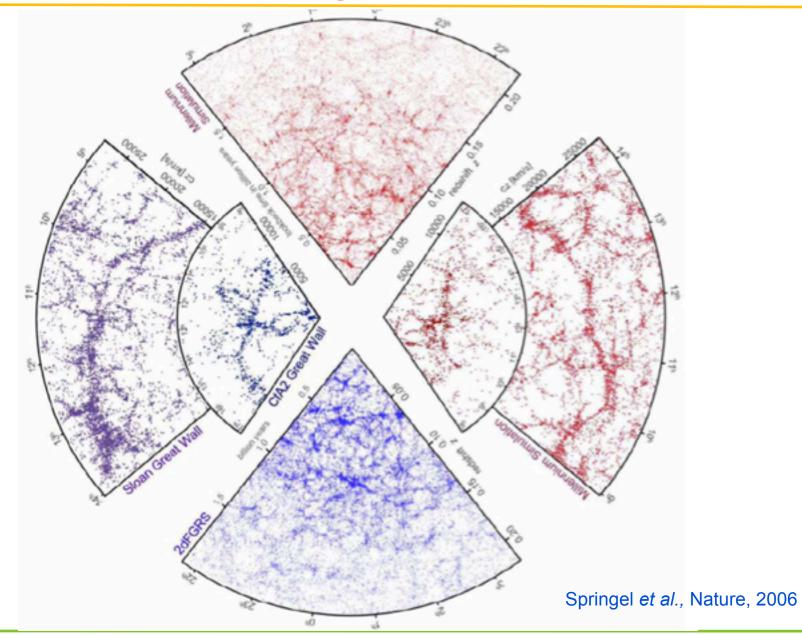


From Gaussian random fields to galaxies: nonlinear dynamics of gravitational instability with Nbody and hydrodynamics codes.





Cosmological simulations



Godunov scheme for hyperbolic systems

The system of conservation laws

$$\partial_t \mathbf{U} + \partial_x \mathbf{F} = 0$$

is discretized using the following integral form:

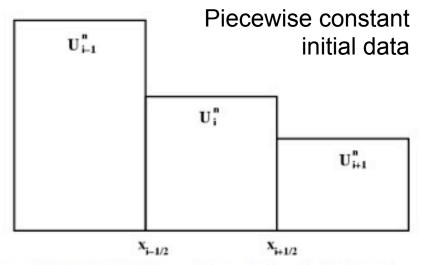
$$\frac{\mathbf{U}_{i}^{n+1} - \mathbf{U}_{i}^{n}}{\Delta t} + \frac{\mathbf{F}_{i+1/2}^{n+1/2} - \mathbf{F}_{i-1/2}^{n+1/2}}{\Delta x} = 0$$

The time average flux function is computed using the self-similar solution of the inter-cell Riemann problem:

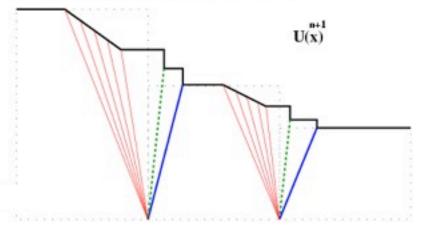
$$\mathbf{U}_{i+1/2}^*(x/t) = \mathcal{RP}\left[\mathbf{U}_i^n, \mathbf{U}_{i+1}^n\right]$$
$$\mathbf{F}_{i+1/2}^{n+1/2} = \mathbf{F}(\mathbf{U}_{i+1/2}^*(0))$$

This defines the Godunov flux:

$$\mathbf{F}_{i+1/2}^{n+1/2} = \mathbf{F}^*(\mathbf{U}_i^n, \mathbf{U}_{i+1}^n)$$



 Godunov, S. K. (1959), A Difference Scheme for Numerical Solution of Discontinuos Solution of Hydrodynamic Equations, *Math. Sbornik*, 47, 271-306, translated US Joint Publ. Res. Service, JPRS 7226, 1969.



Advection: 1 wave, Euler: 3 waves, MHD: 7 waves

Smooth regions of the flow

More efficient to go to higher order.

Spectral methods can show *exponential convergence*.

More flexible approaches: use *ultra-high-order* shockcapturing schemes: 4th order scheme, ENO, WENO, discontinuous Galerkin and discontinuous element methods

Discontinuity in the flow

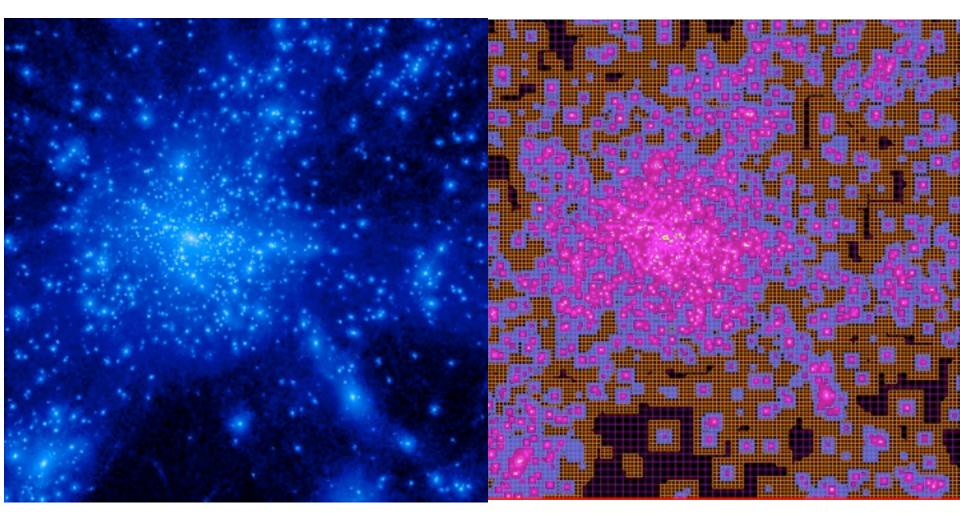
More efficient to refine the mesh, since higher order schemes drop to first order.

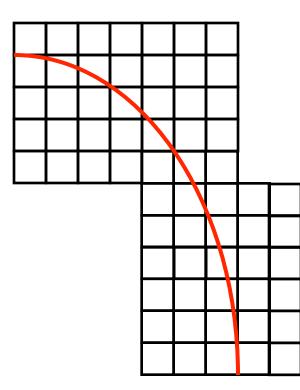
Adaptive Mesh Refinement is the most appealing approach.

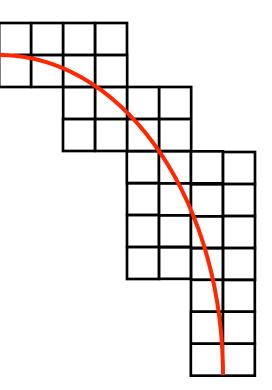
What about the future ?

Combine the 2 approaches. Usually referred to as "*h-p adaptivity*".

Adaptive Mesh Refinement







ENZO: Greg Bryan, Michael Norman, Tom Abel

ART: Andrey Kravtsov, Anatoly Klypin

RAMSES: Romain Teyssier

NIRVANA: Udo Ziegler

AMRVAC: Gabor Thot and Rony Keppens

FLASH: The Flash group (PARAMESH lib)

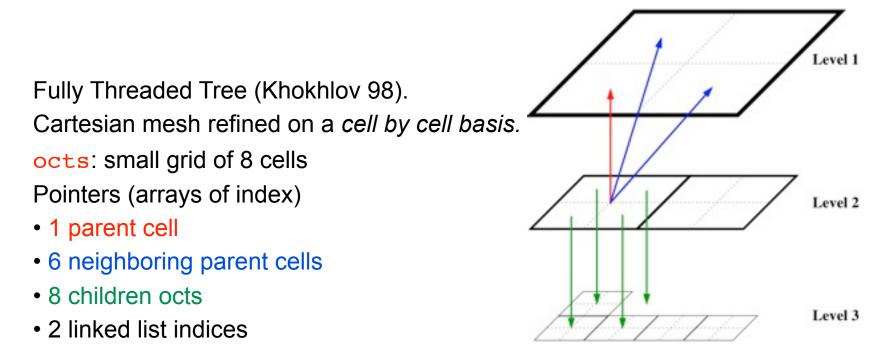
ORION: Richard Klein, Chris McKee, Phil Colella

PLUTO: Andrea Mignone (CHOMBO lib, Phil Colella)

CHARM: Francesco Miniati (CHOMBO lib, Phil Colella)

ASTROBear: Adam Frank

Graded Octree structure



Cell-centered variables are updated level by level using linked lists.

Cost = 2 integer per cell.

Optimize mesh adaptation to complex flow geometries, but CPU overhead compared to unigrid can be as large as 50%.

- 2 type of cell: "leaf" or active cell
 - "split" or inactive cell

Refinement rules for graded octree

Compute the refinement map: flag = 0 or 1

Step 1: mesh consistency

if a split cell contains at least one split or marked cell, then mark the cell with flag = 1 and mark its 26 neighbors

Step 2: physical criteria

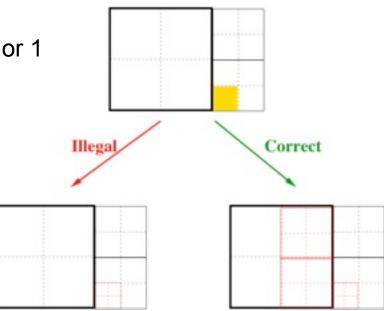
quasi-Lagrangian evolution, Jeans mass

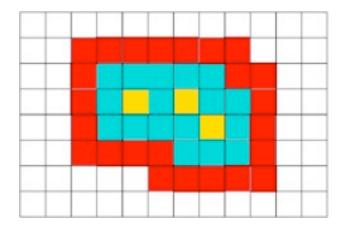
geometrical constraints (zoom)

Truncation errors, density gradients...

Step 3: mesh smoothing

apply a dilatation operator (mathematical morphology) to regions marked for refinement \rightarrow convex hull





Godunov schemes and AMR

Berger & Oliger (84), Berger & Collela (89)

Prolongation (interpolation) to finer levels

- fill buffer cells (boundary conditions)
- create new cells (refinements)

Restriction (averaging) to coarser levels

- destroy old cells (de-refinements)

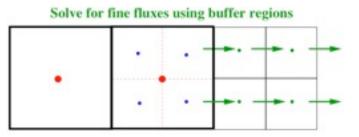
Flux correction at level boundary

$$(\mathbf{F}_{i+1/2,j}^{n+1/2,\ell}) = \frac{(\mathbf{F}_{i+1/2,j-1/4}^{n+1/2,\ell+1}) + (\mathbf{F}_{i+1/2,j+1/4}^{n+1/2,\ell+1})}{2}$$

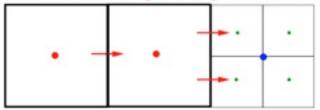
Careful choice of interpolation variables (conservative or not ?)

Several interpolation strategies (with $R^T P = I$):

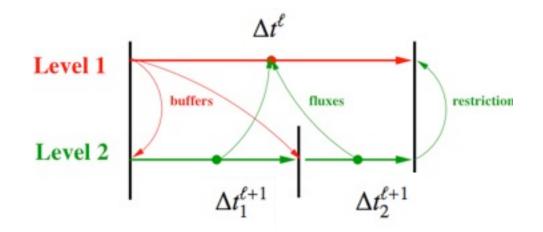
- straight injection
- tri-linear, tri-parabolic reconstruction



Coarse flux: time and space average of fine fluxes



Adaptive Time Stepping

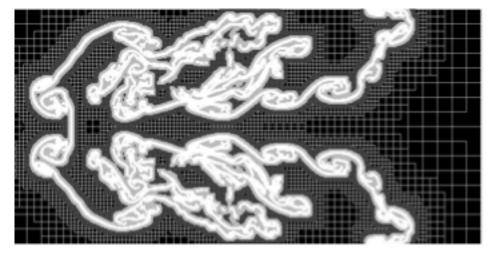


Time integration: single time step or recursive sub-cycling

- froze coarse level during fine level solves (one order of accuracy down !)
- average fluxes in time at coarse fine boundaries

$$(\mathbf{F}_{i+1/2,j}^{n+1/2,\ell}) = \frac{1}{\Delta t_1^{\ell+1} + \Delta t_2^{\ell+1}} \left(\Delta t_1^{\ell+1} \frac{(\mathbf{F}_{i+1/2,j-1/4}^{n+1/4,\ell+1}) + (\mathbf{F}_{i+1/2,j+1/4}^{n+1/4,\ell+1})}{2} + \Delta t_2^{\ell+1} \frac{(\mathbf{F}_{i+1/2,j-1/4}^{n+3/4,\ell+1}) + (\mathbf{F}_{i+1/2,j+1/4}^{n+3/4,\ell+1})}{2} \right)$$

Complex geometry with AMR

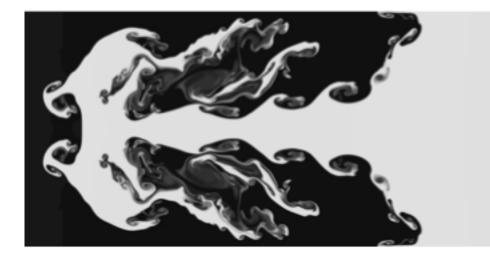


Maximum numerical dissipation occurs at the 2 fluids interface.

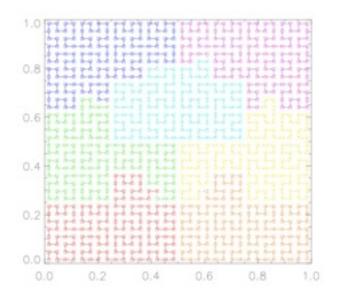
The optimal refinement strategy is based on density gradients.

The number of required cells is directly related to the *fractal exponent n* of the 2D surface.

$$N_{cell} \propto (\Delta x)^{-n}$$



Domain decomposition for parallel computing



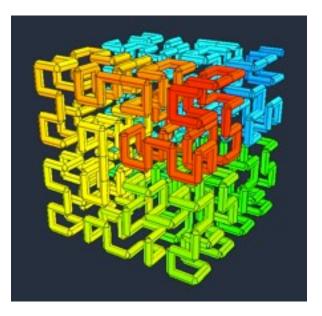
Parallel computing using the MPI library with a domain decomposition based on the *Peano-Hilbert curve*.

Algorithm inspired by gravity solvers (tree codes). Use locally essential tree.

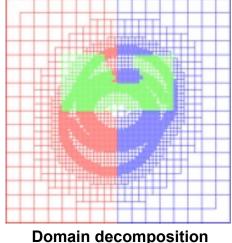
Tested and operational up to 76'000 core. Scaling depends on problem size and complexity.

Salmon, J.K. and Warren, M.S., "Parallel out-of-core methods for N-body simulation", In Eighth SIAM Conference on Parallel Processing for Scientific Computing, SIAM,1997.

Peter MacNeice, Kevin M. Olsonb, Clark Mobarryc, Rosalinda de Fainchteind and Charles Packer, « PARAMESH: A parallel adaptive mesh refinement community toolkit, », 2000, Computer Physics Communications, 126, 3.

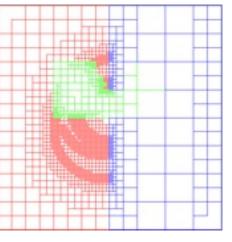


Locally essential trees



over 3 processors

Each processor octree is surrounded by ghost cells (local copy of distant processor octrees) so that the resulting local octree contains all the necessary information.

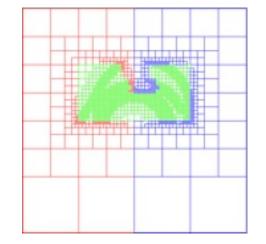


Salmon 90

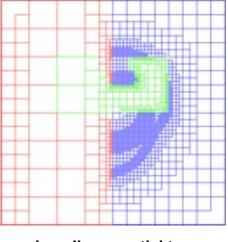
Warren 92

Dubinski 96

Locally essential tree in processor #1



Locally essential tree in processor #2

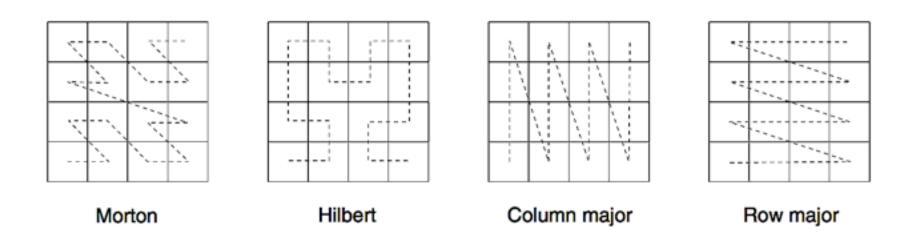


Locally essential tree in processor #3

Dynamic partitioning in RAMSES

Several cell ordering methods:

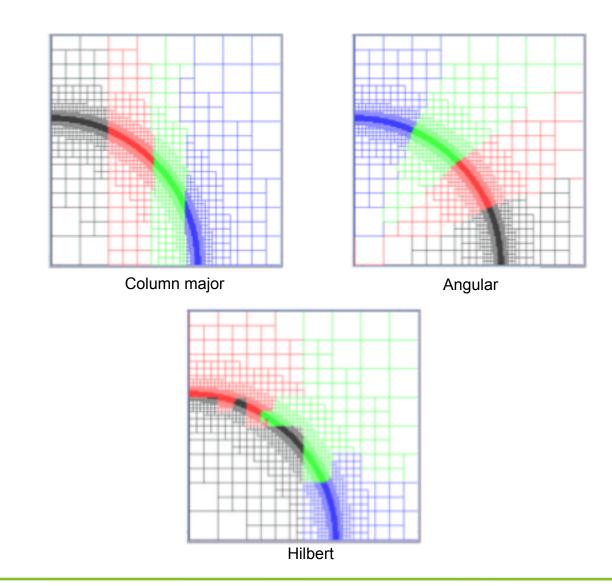
- 1- column, row or plane major
- 2- Hilbert or Morton
- 3- User defined (angular, column+Hilbert...)



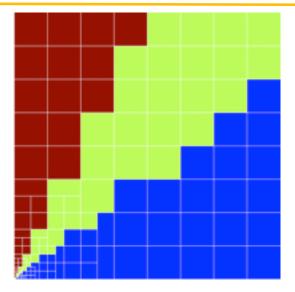
nremap=10

Dynamic partioning is performed every N steps by sorting each cell along chosen ordering and redistributing the mesh across processors. Usually, a good choice is N=10 (10% overhead).

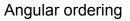
Is there an optimal load balancing strategy ?



Is there an optimal load balancing strategy ?



Recursive bisection

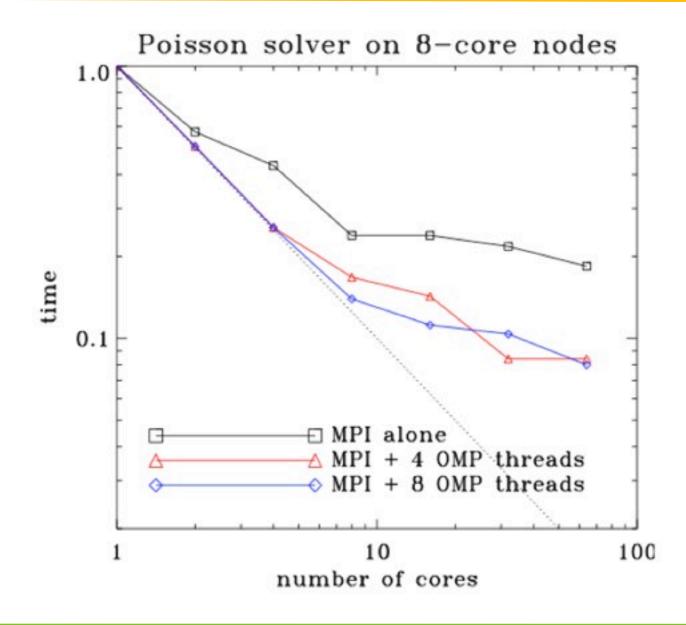


Hilbert ordering

Load balancing issues in RAMSES

Mesh structure	2									
Level 1 has	1 grids	(0,	1,	0,)						
Level 2 has	8 grids	(0,	Ζ,	0,)	1.00					
Level 3 has	64 grids	(0,	7,	0,)	-		-			
Level 4 has	512 grids	(0,	49,	0,)	-					
Level 5 has	4096 grids	(0,	384,	4,)	-		_			
Level 6 has	32768 grids	(0,	3075,	32,)	Je	1				
Level 7 has	262144 grids	(0,	24607,	256,)	time		-			
Level 8 has	21147 grids	(0,	2309,	20,)	eq					
Level 9 has	27690 grids	(0,	2985,	27,)	<u>بة</u> 0.10		-			
Level 10 has	42860 grids	(0,	3948,	41,)	ma					
Level 11 has	85768 grids	(0,	7676,	83,)	Normalized		-			
Level 12 has	213596 grids	(0,	11561,	208,)			-			
Level 13 has	672299 grids	(0,	12783,	656,)	-		1			
Level 14 has	2388883 grids	(0,	21258,	2332,)	-		-			
Level 15 has	2795115 grids	(0,	12304,	2729,)	on 1: comput	o th				
Level 16 has	2850931 grids	(0,	8315,	2784,)	0.01	<u> </u>				
Level 17 has	1662701 grids	(0,	5140,	1623,)	the coarse ¹⁰	eve 100 1000	10000			
Level 18 has	674656 grids	(0,	2859,	658,)		# CPU				
Level 19 has	253539 grids	(0,	1524,	247,)						
Main step= 6735 econs=-1.86E-03 epot=-3.41E-05 ekin= 2.27E-05 eint= 1.57E-08										
Fine step= 728	805 t=-2.65682E+0	0 dt= 3.196	E-05 a= 2.9	955E-01 me	m=13.8% 17.0%	Adaptive time step	s and			
Level= 19 M	New star= 66 To	global domain dec	omn							
==> Level=	19 Step= 4 E									
Fine step= 728	806 t=-2.65679E+0	lead to poor load b	alance.							
Level= 19 New star= 71 Tot= 1604591 Mass= 6.322E-08 Lost= 0.0%										
==> Level=	18 Step= 4 E	rror= 3.163	E-05							
==> Level=	19 Step= 4 E	rror= 2.101	E-05			Ear agab problem	thore is			
Fine step= 728	807 t=-2.65675E+0	For each problem, there is								
Level= 19 M	Level= 19 New star= 88 Tot= 1604679 Mass= 6.323E-08 Lost= 0.0%						a plateau in the parallel			
==> Level= 19 Step= 4 Error= 2.267E-05										
Fine step= 72808 t=-2.65672E+00 dt= 3.845E-05 a= 2.955E-01 mem=13.8% 17.0%						efficiency.				
Level= 19 M	New star= 79 To	ot= 160475	58 Mass= 6	.323E-08 L	ost= 0.0%					

Hybrid OpenMP and MPI approach



Cosmology with RAMSES

Particle-Mesh on AMR grids:

Cloud size equal to the local mesh spacing

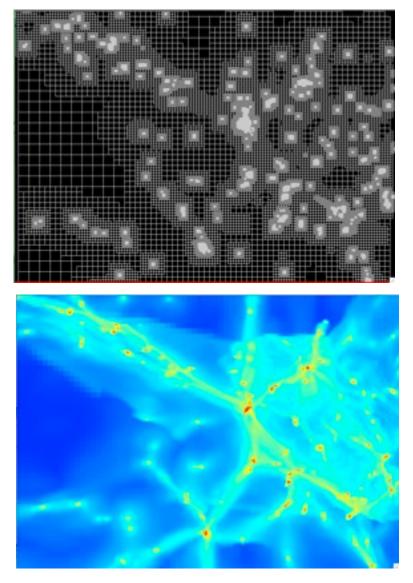
Poisson solver on the AMR grid Multigrid or Conjugate Gradient Interpolation to get Dirichlet boundary conditions (one way interface)

Quasi-Lagrangian mesh evolution:

roughly constant number of particles per cell $n = \frac{\rho_{DM}}{\rho_{DM}} + \frac{\rho_{gas}}{\rho_{gas}} + \frac{\rho_*}{\rho_*}$

$$n = \frac{p_{DM}}{m_{DM}} + \frac{p_{gas}}{m_{gas}} + \frac{p_*}{m_*}$$

Trigger new refinement when n > 10-40 particles. The fractal dimension is close to 1.5 at large scale (filaments) and is less than 1 at small scales (clumps).



Cosmological simulation with RAMSES

&RUN_PARAMS cosmo=.true. pic=.true. poisson=.true. nrestart=0 nremap=10 nsubcycle=1,2 ncontrol=1 /

&OUTPUT_PARAMS noutput=10 aout=0.1,0.2,0.3,0.4,0.5,0.6,0.7,0.8,0.9,1.0 /

```
&INIT_PARAMS
filetype='grafic'
initfile(1)='/scratchdir/grafic_files'
/
```

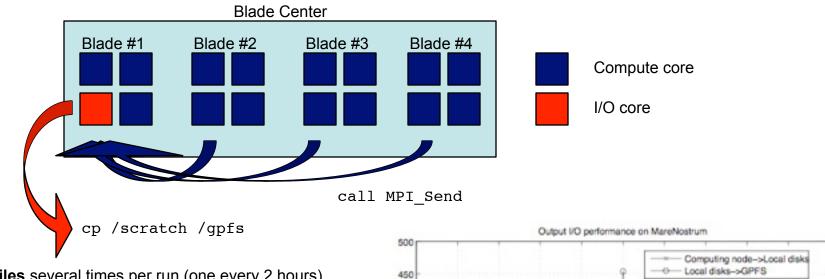
&AMR_PARAMS levelmin=7 levelmax=14 ngridtot=2000000 nparttot=3000000 nexpand=1 / &REFINE_PARAMS

```
m_refine=7*8.
```

Pure N-body case. nrestart controls check-point restarts. nremap controls load-balancing frequency. Choose your output frequency. Initial conditions and cosmological model read in from a set of grafic files. Choose base (coare) grid resolution with levelmin. Choose ultimate refinement resolution with levelmax. Critical value levelmax=21! ngridmax (npartmax) controls the maximum memory available for AMR cells (particles).

Choose refinement strategy (here 8 particles per cell triggers a refinement).

Parallel I/O strategy



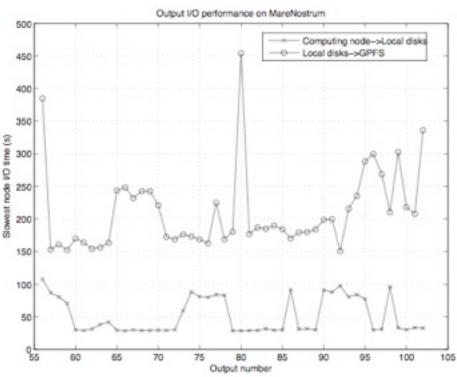
Output files several times per run (one every 2 hours).

Need to minimize computational downtime due to parallel file system contention (competitive disc access).

Compute group: 2048 compute cores : MPI COMM WORLD COMP

I/O group: 1 I/O core per 32 compute cores : MPI COMM WORLD IOGRP (64 dedicated I/O cores)

Compute core: -Write replaced by MPI Send -Wait for compute core to send data I/O core: -Dump data to local scratch disc -Resume computation when all data received -Transfer data from local scratch to GPFS



Very large N body simulations with RAMSES

Billion dollar experiments need support from HPC

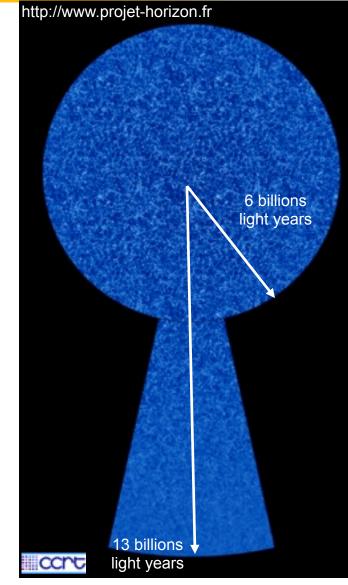




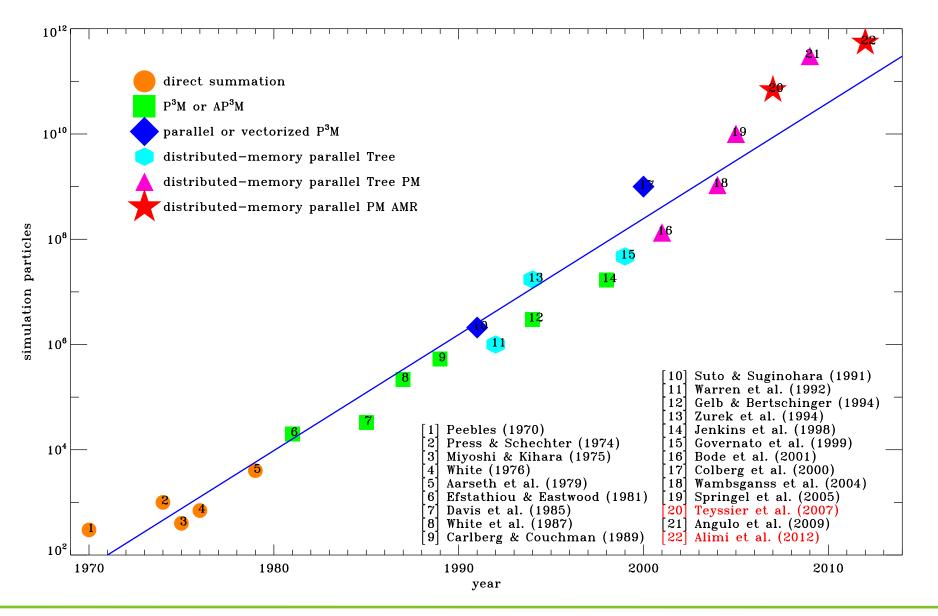
The Horizon simulation (2 Gpc/h) 70 billion dark matter particles and 140 billion AMR cells 6144 core 2.4 GB per core Wall-clock time 2 month performed in 2007 on the CCRT BULL Novascale 3045

The DEUS simulation (21 Gpc/h)

550 billion dark matter particles and 2 trillion AMR cells
76032 core
4 GB per core
Wall-clock time 1 week
performed in 2012 on the CCRT BULL Bullx S6010



Cosmological N body simulations



Conclusions

- RAMSES is a graded octree AMR
- parallel computing using global domain decomposition
- scaling limited by load balancing
- massive N body run up to 76 000 core